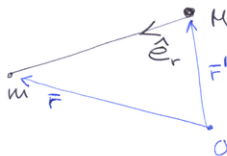


# Þyngdarfræði

Newton: milli tveggja punktmassa er vektor kræftur

$$\vec{F} = -G \frac{mM}{r^2} \hat{e}_r$$



$$G = 6,673 \pm 0,010 \cdot 10^{-11} \text{ Nm}^2/\text{kg}^2$$

Ef hluturinn með massa M er með samfellda dreifingu

$$\rightarrow M = \int dV \rho(r)$$

og

$$\vec{F}(r) = -Gm \int_V \frac{\rho(r')}{|r-r'|^2} \underbrace{\frac{r-r'}{|r-r'|}}_{\text{einingarvígur}} dr'^3$$

$r$  : Atflugandi  
 $r'$  : uppspretta

↑ Þyngdar kræftur

einingarvígur

tyngdensvikt

(2)

$$\bar{g} \equiv \frac{\bar{F}}{m} = -G \frac{M}{r^2} \hat{e}_r$$

eda

$$\bar{g}(\mathbf{r}) = -G \int_V d^3r' \frac{\rho(\mathbf{r}') \hat{e}_{\mathbf{r}-\mathbf{r}'}}{|\mathbf{r}-\mathbf{r}'|^2} \frac{\mathbf{r}-\mathbf{r}'}{|\mathbf{r}-\mathbf{r}'|}$$

tyngdensvikt er sambært vid rafsvik i rafstöðufræði

það er geymt

$$\nabla \times \bar{g} = 0$$

eda

$$\oint_C d\vec{l} \cdot \bar{g} = 0$$

Til gæmans: Hve langt nær sambærður á þyngdarfræði og rafsegulfræði? (3)

Maxwell

$$\nabla \cdot \vec{E} = \frac{\rho}{\epsilon_0}$$

$$\nabla \cdot \vec{B} = 0$$

$$\nabla \times \vec{E} = -\frac{\partial \vec{B}}{\partial t}$$

$$\nabla \times \vec{H} = \vec{J} + \frac{\partial \vec{D}}{\partial t}$$

$$\vec{D} = \epsilon \vec{E}$$

$$\vec{H} = \frac{1}{\mu} \vec{B}$$

$$\vec{E} = \varphi(\vec{E} + \vec{v} \times \vec{B})$$

Almennu sviðsjöfnur Einsteins →  
gæðar línulegar

$$\nabla \cdot \vec{g} = -4\pi G \rho$$

$$\vec{F}_g = m(\vec{g} + 4\vec{v} \times \vec{b})$$

$$\nabla \cdot \vec{b} = 0$$

$$\nabla \times \vec{g} = -\frac{\partial \vec{b}}{\partial t}$$

$$\nabla \times \vec{b} = -\frac{4\pi G}{c^2} \vec{J}_g + \frac{1}{c^2} \frac{\partial \vec{g}}{\partial t}$$

↑ segulklefi þyngdar sviðs

GPS

þyngdarbylgjur?

En við höldum okkur  
við stöðu fræðna

# Þynglarerfi

Samantektir við rafstöðufræðina, geyminn kraftur.....

Til er erfi  $p.a.$

$$\vec{g} = -\nabla\Phi$$

Vidd

$$[\vec{g}] = [\vec{a}] = \frac{L}{T^2} \quad \rightarrow \quad [\Phi] = \frac{L^2}{T^2} \quad ([\nabla] = \frac{1}{L})$$

$$\rightarrow [m\Phi] = M \frac{L^2}{T^2} : \text{vidd orku}$$

Fyrir punktmassa

$$\vec{g} = GM \frac{\hat{e}_r}{r^2}$$

ef punktmassinn er í  
miðju kúta kerfisins

$$\vec{g} = -\nabla\Phi \quad \rightarrow \quad \Phi = -GM \frac{1}{r}$$

← sleppum heildunarfata  
Munum ávallt kosta önnur  
 $\vec{a} \quad \Delta\Phi$

$$\Phi(\vec{r}) = -G \int_V d^3r' \frac{\rho(\vec{r}')}{|\vec{r} - \vec{r}'|} \quad \text{fyrir samfellda massahefningu}$$

Vinna ytri krafts / massa til að flytja hann um  $d\vec{r}$

$$dW' = -\vec{g} \cdot d\vec{r} = (\nabla \Phi) \cdot d\vec{r} = \sum_i \frac{\partial \Phi}{\partial x_i} dx_i = d\Phi$$

Stöðuorka

$$U = m\Phi$$

$$(U = qV)$$

$$\rightarrow \vec{F} = -\nabla U$$

# Lögmál Gauss og jafna Poissons'

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Í bókinni er litið út lögmál Gauss

$$\boxed{\nabla \cdot \vec{g} = -4\pi G \rho} \quad \text{eða} \quad \boxed{\oint_S \vec{g} \cdot d\vec{s} = -4\pi G M}$$

Þegar var komið  $\vec{g} = -\nabla \Phi$

$$\rightarrow \nabla \cdot (\nabla \Phi) = 4\pi G \rho \quad \text{eða} \quad \boxed{\nabla^2 \Phi = 4\pi G \rho}$$

Jafna Poissons gefur  $\Phi$  ef  $\rho$  er þekkt

skalar jafna og þú oft mjög handleg í stað vigrjöfnu, en leusnar ætíð þú þetta ætíð og leyst  $G:IV \dots$

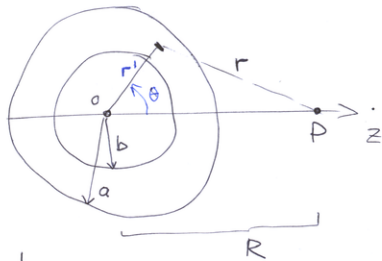
# Demi : Syngbr matter um agi keturkel

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Notum

$$\Phi(r) = -G \int_V dr' \frac{\rho(r')}{|r - r'|}$$

Setjann miðu kúta kerfis í miða skel  
og notum  $\phi$  samhverfu



$$\Phi(R, 0, 0) = -2\pi\rho G \int_b^a r'^2 dr' \int_0^\pi \sin\theta d\theta \frac{1}{r}$$

$$r^2 = r'^2 + R^2 - 2r'R \cos\theta$$

$$\rightarrow 2r dr = 2r'R \sin\theta d\theta \rightarrow \sin\theta d\theta \frac{1}{r} = \frac{dr}{r'R}$$

$$\rightarrow \Phi(R, 0, 0) = -\frac{2\pi\rho G}{R} \int_b^a r' dr' \int_{r_{\min}}^{r_{\max}} dr$$

utan stjörur  $R > a$

$$\Phi(R, 0, 0) = - \frac{2\pi\rho G}{R} \int_b^a r' dr' \int_{R-r'}^{R+r'} dr$$

$(R+r' - R+r') = 2r'$

$$= - \frac{4\pi\rho G}{R} \int_b^a r'^2 dr' = - \frac{4}{3} \frac{\pi\rho G}{R} (a^3 - b^3)$$

Massi stjörur  $M = \frac{4\pi}{3} (a^3 - b^3) \rho$

$$\Phi(R, 0, 0) = - \frac{GM}{R}$$

$R > a$

utan stjörur er myndstæðan öháð massahefningu svo fremi hún sé öháð  $\phi$  og  $\theta$

↑ má sjá með lögmáli Gauss

Stöðsetning þ m.t.t.  $\phi$  og  $\theta$  á língu að breyta, valdi til þess.



Innan steljar:  $R < b$

$$\Phi(R, 0, 0) = - \frac{2\pi\rho G}{R} \int_b^a r' dr' \int_{r'_{\min}}^{r'_{\max}} dr$$

$r'_{\max} = r' + R$   
 $r'_{\min} = r' - R$   
 $(r' + R - r' + R)$

$$= - 4\pi\rho G \int_b^a r' dr' = - 2\pi\rho G (a^2 - b^2)$$

fasti öndæð  
R innan steljar

Inni í stel:  $b < R < a$

$$\Phi(R, 0, 0) = - \frac{2\pi\rho G}{R} \int_b^a r' dr' \left[ \int_R^{r'+R} dr + \int_{R-r'}^R dr \right]$$

$$= - 4\pi\rho G \left\{ \frac{a^2}{2} - \frac{b^2}{3R} - \frac{R^2}{6} \right\}$$

# Reynnum stíran æðfrætt

Notum kúluhit (r,  $\phi$ ,  $\theta$ ) með mætti  $\rho$  í miðju skeljar (10)

Gættum notað Gauss og reiknað þyngdarstöð  $\bar{g}$ , en við höfum

$$\nabla^2 \Phi = 4\pi G \rho$$

$\rho$  er fasti óháður  $\phi$  og  $\theta$ , skelin fellur að kúluhitum  
 $\rightarrow \Phi$  er líka óháð  $\phi$  og  $\theta$ . Í kúluhitum er jafnan

því  $\frac{1}{r^2} \frac{\partial}{\partial r} \left\{ r^2 \frac{\partial \Phi(r)}{\partial r} \right\} = 4\pi G \rho$

Ein breyta eftir skritum því

$$\frac{d}{dr} \left\{ r^2 \frac{d}{dr} \Phi(r) \right\} = 4\pi r^2 G \rho$$

← heildum ákvaðid  
á þremur svæðum

→  $\rho \neq 0$  aðeins á (II)



Kreppumst samfelldni  $\Phi$

(I) :

$$\frac{d}{dr} \left\{ r^2 \frac{d}{dr} \Phi(r) \right\} = 0$$

$$\underline{r < b}$$

(11)

$$\rightarrow r^2 \frac{d}{dr} \Phi(r) = C_1 \quad : \text{konst.}$$

$$\rightarrow \frac{d}{dr} \Phi(r) = \frac{C_1}{r^2} \quad \rightarrow \Phi_I(r) = -\frac{C_1}{r} + C_2 = C_2$$

↙  
 andere mögliche  
 Lösung für  $r > a$   
 $\rightarrow C_1 = 0$   
 in I

(III)

samekonver Lösung für  $r > a$ , in  $\infty$  Bereich  
 lösen

$$\Phi_{III}(r) = -\frac{C_3}{r} + C_4$$

(II)

$$\frac{d}{dr} \left\{ r^2 \frac{d}{dr} \Phi(r) \right\} = 4\pi r^2 G \rho$$

$$b < r < a$$

$$\rightarrow r^2 \frac{d}{dr} \Phi(r) = \frac{4\pi}{3} r^3 G \rho + C_5 \quad \frac{d}{dr} \Phi(r) = \frac{4\pi}{3} r G \rho + \frac{C_5}{r^2}$$

$$\rightarrow \Phi_{II}(r) = \frac{4\pi}{6} r^2 G \rho - \frac{C_5}{r} + C_6$$

Skeytum saman i  $r=a$ ,  $r=b$

Veljum að  $\Phi_{III}(r \rightarrow \infty) = 0 \rightarrow C_4 = 0$

$r=a$   $\Phi_{III}(a) = \Phi_{II}(a) \rightarrow$

$$-\frac{C_3}{a} = \frac{4\pi a^2 G \rho}{6} - \frac{C_5}{a} + C_6$$
  
$$C_2 = \frac{4\pi b^2 G \rho}{6} - \frac{C_5}{b} + C_6$$

$r=b$   $\Phi_I(b) = \Phi_{II}(b) \rightarrow$

4 fastar, tveir jöfnur, þessum líka samfelldu  $\vec{g} = -\nabla\Phi$   
 $\vec{g}$  er aðeins með útpátt (radial)  $\vec{g} \cdot \hat{e}_r = -\frac{\partial}{\partial r}\Phi$

$r=a$   $\Phi'_I(a) = \Phi'_{II}(a) \rightarrow$

$$\frac{C_3}{a^2} = \frac{4\pi a}{3} G \rho + \frac{C_5}{a^2}$$
  
$$0 = \frac{4\pi b}{3} G \rho + \frac{C_5}{b^2}$$

$r=b$   $\Phi'_I(b) = \Phi'_{II}(b) \rightarrow$

$$-C_3 + C_5 - aC_6 = \frac{4\pi a^3 G \rho}{6}$$

$$C_5 = -\frac{4\pi b^3 G \rho}{3}$$

$$C_2 + \frac{C_5}{b} - C_6 = \frac{4\pi b^2 G \rho}{6}$$

$$C_3 = \frac{4\pi G \rho}{3} \{a^3 - b^3\} = M$$

$$C_3 - C_5 = \frac{4\pi a^3 G \rho}{3}$$

$$-C_5 = \frac{4\pi b^3 G \rho}{3}$$

$$-aC_6 = \frac{4\pi a^3 G \rho}{6} + C_3 - C_5 = \frac{4\pi G \rho}{6} \{a^3 + 2a^3\} = 2\pi G \rho a^3$$

$$\rightarrow C_6 = -2\pi G \rho a^2$$

$$C_2 = \frac{4\pi b^2 G \rho}{6} + C_6 - \frac{C_5}{b} = \frac{4\pi G \rho}{6} \left\{ b^2 - \frac{6}{2} a^2 + 2b^2 \right\} = 2\pi G \rho \{ b^2 - a^2 \}$$

pu erlausnin

(14)

$$\left\{ \begin{array}{l} \Phi_{\text{I}}(r) = -2\pi G\rho \{a^2 - b^2\} \quad \underline{r < b} \\ \Phi_{\text{II}}(r) = -4\pi G\rho \left\{ -\frac{r^2}{6} - \frac{b^3}{3r} + \frac{a^2}{2} \right\} \quad \underline{b < r < a} \\ \Phi_{\text{III}}(r) = -\frac{4\pi G\rho}{3r} \{a^3 - b^3\} = -\frac{GM}{r} \quad \underline{r > a} \end{array} \right.$$

og því

$$\bar{g}_{\text{I}}(r) = 0$$

$$\bar{g}_{\text{II}}(r) = \frac{4\pi G\rho}{3} \left\{ \frac{b^3}{r^2} - r \right\} \hat{e}_r$$

$$g_{\text{III}}(r) = -\frac{GM}{r^2} \hat{e}_r$$

